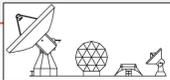


Yebe Observatory: Future Core Site and Laser Ranging Station Status

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YEBES OBSERVATORY

Yebe Observatory, founded in the late 70s, is located in the center of Spain (80km from Madrid), a strategic place in the limit of the European Tectonic Plate. Thanks to the RAEGE Project (Atlantic Network of Geodynamical and Space Stations) and the European Regional Development Fund (ERDF) the Observatory will become in the next years the first Core Site in the Iberian Peninsula. Currently, there are two space geodesy techniques present in the Observatory (VLBI) and GNSS) in addition to the Gravimetry laboratory (absolute gravimeters and superconducting relative gravimeter), the local tie network, time and frequency system and other facilities.

To fulfill a Fundamental Geodetic Station, following the requirements of the GGOS Project, just a Laser Ranging System is necessary to be added. The project for building a Laser Ranging Station at Yebe Observatory has just begun. The station applications (Satellite Laser Ranging and Space Debris tracking) and main specifications are already established. First contacts with the sector companies and institutions have been carried out and the funds are available for building the complete system, allowing Yebe Laser Ranging Station (YLARA) to be in operation by the end of 2020. Fulfilling the demanded ILRS requirements would allow the station to be included into the ILRS network.

OBSERVATORY OVERVIEW



40 m RT. Geodesy and astronomy applications. From 2 to 100 GHz.



State of art laboratories: cryogenic receivers, low noise amplifiers, up/down converters, mechanical workshops...



Anechoic Chamber Planar and spherical near field systems



- Madrid (80 km)
- Main building, laboratories and workshops
- Meteo station
- 14 m RT
- Gravimetry
- Astrograph
- Outreach pavillion
- GNSS Receivers (2)
- Local tie (20 pillars)
- 40 m Radio Telescope
- Hydrogen masers
- Anechoic Chamber
- Solar Tower
- 13,2 m Radio Telescope

THE RAEGE PROJECT

Establishment of an Spanish-Portuguese Network of Geodynamical and Space Stations (RAEGE) by the installation and operation of **four fundamental geodetic stations provided with** radio telescopes with VGOS specifications in Yebe (1), Canary Islands (1) and Azores Islands (2).



Initial equipment to be installed at each RAEGE station:

- Geodetic VGOS radio telescope:
 - Diameter 13.2 m, frequency ≥ 40 GHz
- Gravimeter
- Permanent GNSS station
- Satellite Laser Ranging (at Yebe Observatory)

Radio Telescope characteristics

- Maximum AZ/EL velocity $12^\circ/\text{sec}$ and $6^\circ/\text{sec}$ (acceleration $3^\circ/\text{sec}^2$)
- Upper operational frequency 40 GHz (200 microns rms surface accuracy).
- Possible upgraded to 100 GHz. Path length error < 0.26 mm.
- Physical range in elevation $0 - 100^\circ$, azimuth $540^\circ (+270 \text{ degrees})$
- Overall pointing precision < 16 arcsec
- Power consumption < 170 kW



CORE SITE STATUS - AVAILABLE TECHNIQUES

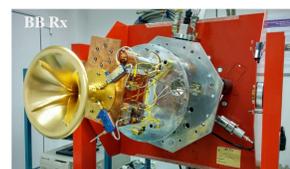
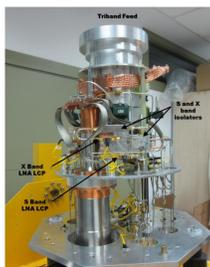
Gravimetry pavillion



- Seven pillars for instrument installation and intercomparisons
- Two absolute gravimeters (A10 & FG5)
- OSG Superconducting gravimeter (Feb - 2012)
- Participation in IGETS - International Geodynamics and Earth tide Service



13.2 m VGOS Radio Telescope



LNAs
Ka band (25-35 GHz)
X band (4-12 GHz)



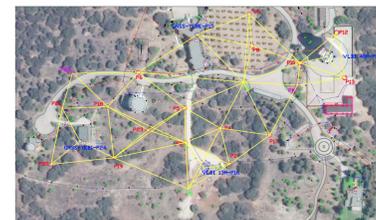
Up/down converters

Yebe Observatory Developments for VGOS

- **Triband Receiver**
 - S band: 2.2-2.6 GHz, X band: 7.5-9.0 GHz, Ka band: 28-33 GHz
 - Installed in February 2014 for radio telescope commissioning
 - Other Triband systems: Japan and Santa María in Azores
 - Future installation at Ny-Ålesund Observatory (NMA)
- **Broadband Receiver (from 2 to 14 GHz)**
 - Dual lineal polarization, noise temperature under 25 K
 - Using a quadruple-ridged flared horn (QRFH) from Caltech
 - Installed in February 2016 for VGOS Broadband observations
- Low Noise Amplifiers: S, X and Ka bands and broadband for VGOS
- Frequency up/down converters
- Phase and gain calibration modules

First Transatlantic VGOS Fringes (june 2016): Yebe antenna was one of the radio telescopes involved in the observations with Kokee Park (Hawaii), GGAO (Maryland), Westford (Massachusetts), Wettzell (Germany) using broadband systems.

Local tie



Studies to define the best network configuration allow us to get the local tie with an accuracy below 1 mm. The complete network is composed by 24 vertex including on it the radio telescopes and the GNSS antennas.



Pillars are made of concrete and iron and compose by a 30 cm diameter cylinder inside a protector tube



GNSS Receivers



Spanish Network ERGNSS



EUREF Permanent Network

Main Specifications

Observations to satellites equipped with retro-reflector

- Satellites observations from 200 - 42000 km
- Pulsed Solid State Laser
 - Repetition rate ≥ 1000 Hz
 - Pulsed width < 25 ps
 - Wavelength 532 nm (Nd:YAG - Nd:Van)
 - Energy 0.5 - 2 mJ
- Biaxial telescope, AZ-EL mount
 - Receiving system 50-60 cm
 - Transmitting system 10 cm
 - Pointing accuracy 5-7"
 - High slew rate
- Detector CSPAD (QE 20%)

- Range Gate Width 100-400 ns
- Frequency Standard: Hydrogen maser
- Event timer
- Highly automation system
- Aircraft safety system compatible with other geodetic systems in the observatory

Space debris observation capacity

- To be determined:
 - Monostatic/bistatic observation
 - Laser system characteristics (power, color, repetition rate, etc.)
 - Specific software

- **Yebe Observatory is one of the Singular Scientific and Technological Infrastructures (ICTS)** in Spain, the only one in the Castilla-La Mancha Region. The ICTS are facilities, resources, equipment and services, unique in its kind, and dedicated to cutting edge high quality research and development, to promote transfer, exchange and preservation of knowledge, technology and innovation.

⇒ **ERDF 2014-2020 Funds - Ministry of Economy and Competitiveness of Spain** (FICTS1420-11-12)

- The Economic and Technical ERDF Memory establishes the next preliminary work schedule for the construction of the YLARA system:

Work Schedule	2016	2017	2018	2019	2020
SLR building and infrastructure		X	X	X	X
Telescope and dome subsystems		X	X	X	X
Optic Subsystem		X	X	X	X
Measurement subsystem		X	X	X	X
Software Package and security system		X	X	X	X
System Engineering and Integration		X	X	X	X
Project Management and Quality Assurance	X	X	X	X	X
Promotion	X	X	X	X	X